

The Heisenberg model for high energy
nucleon-nucleon
scattering and AdS/CFT

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SILAFEA, Antigua Guatemala,
November 2016

based on: (K. Kang and H.N. PLB 2005 and) H.N., J. Sonnenschein PRD
2015, H.N. JHEP 2016

Summary:

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0. Introduction

- At $s \rightarrow \infty$ (QCD) collision cross section is bound by unitarity (Froissart bound)

$$\sigma_{\text{tot}}(s) \leq C \ln^2 \frac{s}{s_0}, \quad C \leq \frac{\pi}{m_\pi^2}$$

- But 1952 Heisenberg model saturates bound (in QCD: hard!)
- Use AdS/CFT to look again at the Heisenberg model
- Asymptotic regime \leftrightarrow scattering creating BHs in IR.
- Asymptotic regime: nucleon-nucleon (e.g. pp) collision at LHC/
heavy ion collisions at RHIC/LHC.

- We will extend extend the Heisenberg model, show that it is quite unique
- We will extend to corrections to $\sigma_{\text{tot}}(s)$, as well as $\sigma_{\text{elastic}}(s)$.
- We will suggest an extension to be used for the sQGP hydrodynamics
- We will obtain "DBI scalar viscous hydrodynamics" for sQGP
→ restricted hydro model.

1. AdS/CFT and high energy scattering

- String theory: all fundamental particles (matter and gauge bosons) are actually strings, living in $D = 10$ spacetime dimensions, with 16 or 32 supercharges.
- KK compactification to 3+1 dimensions + breaking susy to $\mathcal{N} = 1 \Rightarrow$ should get MSSM (susy version of Standard Model)
- But instead: AdS/CFT: model strongly coupled field theories via *duality* to string theory in some *curved* background (**not** real world)
- Standard example: $SU(N)$ $\mathcal{N} = 4$ SYM in 3+1 dimensions \leftrightarrow string theory in $AdS_5 \times S^5$ background, for $N \rightarrow \infty$, $g^2 N =$ fixed, large.

- But interested in high energy scattering in QCD \rightarrow with Λ_{QCD} .

- Polchinski-Strassler (2001): "hard-wall model"

$$ds^2 = \frac{r^2}{R^2} d\vec{x}^2 + \frac{R^2}{r^2} dr^2 + R^2 ds_X^2, \quad r \leq r_{\min} = R^2 \Lambda_{\text{QCD}}.$$

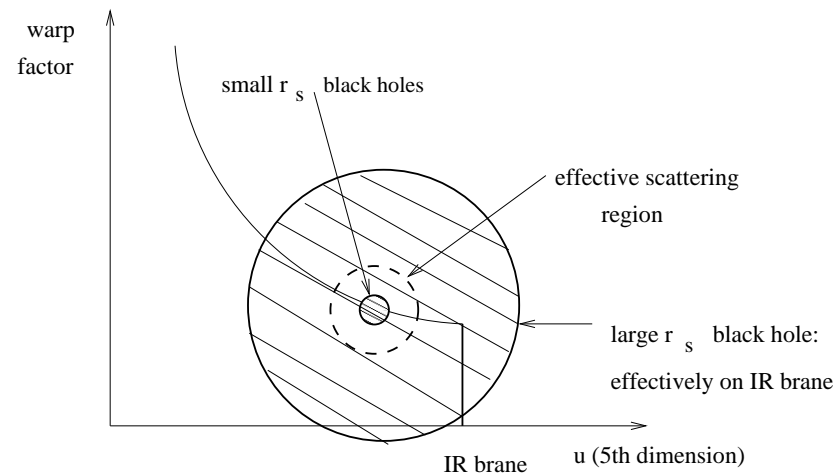
- Momenta rescaled: $\tilde{p}_{(AdS)} = \frac{R}{r} p_{(QCD)}$.

- "Glueball" wavefunction $e^{ip \cdot x} \leftrightarrow$ AdS wavefunction $e^{ip \cdot x} \psi(r, \Omega)$.

$$\mathcal{A}_{\text{QCD}}(p_i) = \int dr d^5 \Omega \sqrt{g} \mathcal{A}_{\text{string}}(\tilde{p}_i) \prod_i \psi_i$$

- From $r \geq r_{\min} \Rightarrow \sqrt{\alpha'} \tilde{p} \leq \sqrt{\hat{\alpha}''} p$, $\hat{\alpha}' = \sqrt{g_{YM}^2 N} \Lambda_{\text{QCD}}^{-2} =$ QCD string tension.

- Find various regimes in QCD (HN, 2004 and 2005). Max behaviour: Froissart saturation \leftrightarrow create black holes on the IR brane (cut-off).



Scattering in the cut-off AdS gravity dual. Small black holes are effectively produced in a scattering region and don't feel the IR brane. Asymptotically large black holes live effectively on the IR brane and have negligible quantum fluctuations in the 5th dimension.

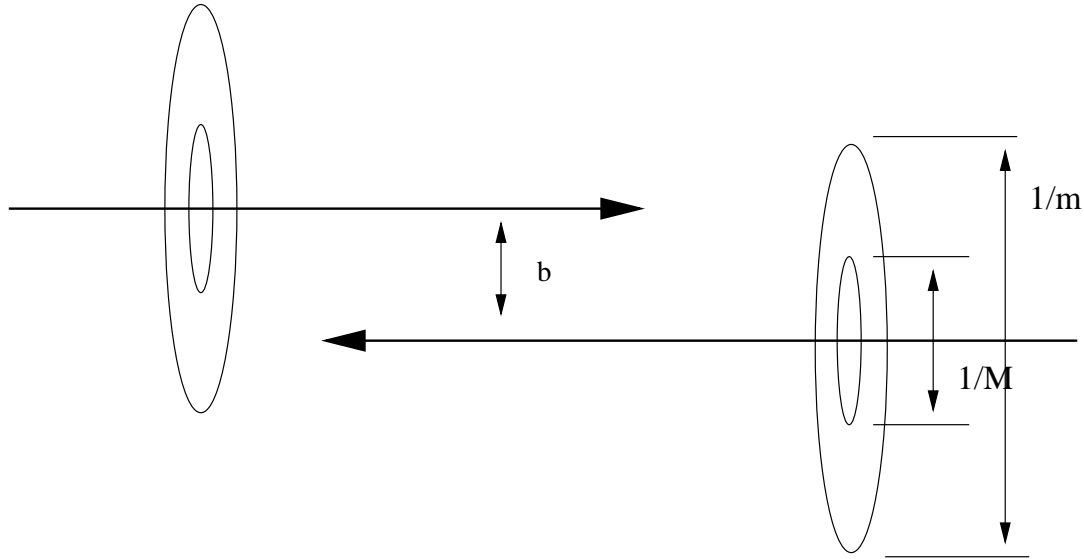
- In the IR limit, BH \simeq classical. Corresponds to sQGP fireball (HN, 2005): \exists classical field theory description?
- Yes: Heisenberg's 1952 model.
- High energy gravitational scattering at $\sqrt{s} \sim M_{\text{Pl}}$: gravitational shockwave scattering. (K. Kang and HN, 2005, HN 2005)

2. The original Heisenberg's model

- Ideas: at $s \rightarrow \infty$, QCD scattering dominated by pion (lightest particle) *classical* field: large number of pions produced.
- But pion action is neither free, nor potential oriented (will see: it doesn't work) \rightarrow is nonlinear in derivatives: massive DBI.

$$\mathcal{L} = l^{-4} \left[1 - \sqrt{1 + l^4 [(\partial_\mu \phi)^2 + m^2 \phi^2]} \right].$$

- Asymptotic: scattering of pion field shockwaves (Lorentz contraction: pancakes \rightarrow shockwaves for both source and field)



Hadron scattering in the center of mass frame. M = hadron mass, m = pion mass. Also, A-S shockwave scattering on the IR brane. M = dual particle size. m =KK graviton mass (gravitational field in 5d, with given boundary conditions)

- Find $\sigma_{\text{tot}}(s)$ from a combination of classical and quantum: average emitted energy per pion $\langle k_0 \rangle = \mathcal{E}/n$ from classical scalar field + quantization (momentum $k \equiv$ pion energy k_0).

- Then, wavefunction at large r , $\psi(r) \sim e^{-m_\pi r}$ and energy loss $\mathcal{E} \propto \sqrt{s}$ and $\propto \psi(b) \sim e^{-m_\pi b}$. At b_{max} , emit *one* pion \Rightarrow

$$(\mathcal{E} =) \langle k_0 \rangle = \frac{\mathcal{E}}{n} = \sqrt{s} e^{-m_\pi b_{\text{max}}}$$

- Then, cross section = *classical* one,

$$\sigma_{\text{tot}}(s) = \pi b_{\text{max}}^2 = \frac{\pi}{m_\pi^2} \log^2 \frac{\sqrt{s}}{\langle s \rangle}.$$

- But only for DBI we find $\langle k_0 \rangle \simeq \text{const.}$ For free or $V = \sum_p \lambda_p \phi^p$, $\langle k_0 \rangle \propto \sqrt{s}$.

- Shockwave limit: $\phi = \phi(s)$, where $s = t^2 - x^2 \rightarrow$ boost invariant solution needed to move on $x^\pm = 0 \rightarrow$ restricts $\phi(x^+, x^-)$ to $\phi(x^+ x^-) = \phi(s)$.

- On $\phi = \phi(s)$, $(\partial_\mu \phi)^2 = -4s(d\phi/ds)^2$, and

$$\mathcal{L} = l^{-4} \left[1 - \sqrt{1 + l^4 \left(-4s \left(\frac{d\phi}{ds} \right)^2 + m^2 \phi^2 \right)} \right]$$

- Solution

$$\phi = \frac{\sqrt{s}}{l^2} (1 + a s m^2 + \dots), \quad 0 \leq s \ll 1/m^2$$

- But need $(\partial_\mu \phi)^2$ finite at the shock $\Rightarrow \phi \sim A\sqrt{s}$ as $s \rightarrow 0$. Free theory: $\phi = A \log s/s_0 \Rightarrow (\partial_\mu \phi)^2 = -4A^2/s \rightarrow \infty$.

- Energy density:

$$\mathcal{H} = \pi\dot{\phi} - \mathcal{L} = \frac{l^{-4} + (\nabla\phi)^2 + m^2\phi^2}{\sqrt{1 + l^4[(\partial_\mu\phi)^2 + m^2\phi^2]}} - l^{-4}$$

- Perturbations: Square root finite, instead of zero. $\rightarrow \exists$ cut-off
- Find

$$\phi - l^{-2}i\frac{|t|}{|k|} \propto |t|^{1/2}|k|^{-3/2}e^{-i|k||t|} \Rightarrow$$

$$\frac{dE}{dk} \propto k^2\phi(k)^2 \sim \frac{\text{const.}}{k} \quad \text{for } m \leq k \leq m\gamma$$

and $k = k_0$ (pion energy), $E = \mathcal{E}$. One finally finds

$$\langle k_0 \rangle \equiv \frac{\mathcal{E}}{n} = m \frac{\ln \gamma}{1 - \frac{1}{\gamma}} \simeq \text{const.}$$

- But for $V = \sum_p \lambda_p \phi^p \Rightarrow$ find $d\mathcal{E}/dk_0 = B = \text{const.} \Rightarrow$

$$\langle k_0 \rangle \propto \frac{\sqrt{\tilde{s}}}{\ln \sqrt{\tilde{s}}} \Rightarrow \text{not good}$$

3. AdS/CFT description of saturation of Froissart bound

- High enough energy \Rightarrow scattering of Aichelburg-Sexl gravitational shockwaves (BHs boosted to $v = c$)

$$ds^2 = 2dx^+ dx^- + (dx^+)^2 \Phi(x^i) \delta(x^+) + d\vec{x}^2$$

$$\Delta_{D-2} \Phi(x^i) = -16\pi G p \delta^{D-2}(x^i)$$

- When

$$R > M_{\text{Pl}} = g_s^{-1/4} \alpha'^{-1/2} = g_s^{-1/4} (R^2 / \sqrt{g_s N})^{-1/2} = N^{1/4} R$$

$$\Leftrightarrow N^{1/4} \Lambda_{\text{QCD}} \equiv \hat{M}_{\text{Pl}}$$

2 shockwaves collide, producing BHs (Giddings, 2002)

- When BH reaches AdS size, $E \sim M_{\text{Pl}}^8 R^7 \Rightarrow$ at $E_R = M_{\text{Pl}} (M_{\text{Pl}} R)^7$.
- At $E \sim E_F \Leftrightarrow E = \hat{E}_F$, depending on details of IR of gravity dual, BH reaches cut-off: max. behaviour: Froissart. (K. Kang, H.N., 2004)

- Simple argument (Giddings) linearized perturbation *on IR brane*.

$$h_{00,lin} \sim G_4 \sqrt{s} \frac{e^{-M_1 r}}{r} \sim 1; \quad G_4^{-1} = M_P^3 R$$

$$\Rightarrow r = r_H \Rightarrow \sigma \sim \pi r_H^2 \sim \frac{\pi}{M_1^2} \ln^2(\sqrt{s} G_4 M_1)$$

- Here $M_1 = j_{1,1}/R =$ lightest KK mode on IR brane \leftrightarrow glueball. For pion (\sim fluctuation of D-brane position L)

$$\frac{\delta L}{L}|_{lin} \sim G_4 \sqrt{s} (M_L R) \frac{e^{-M_1 r}}{r} \sim 1 \Rightarrow \sigma_{\text{QCD}} \sim \frac{\pi}{M_L^2} \ln^2(\sqrt{s} G_4 M_L R M_1)$$

- But, more precise: use *shockwaves* on IR brane. As $r \rightarrow \infty$, at $y = 0$ (IR brane), profile is

$$\Phi(r, y = 0) \simeq R_s \sqrt{\frac{2\pi R}{r}} C_1 e^{-M_1 r};$$

$$C_1 = \frac{j_{1,1}^{-1/2} J_2(j_{1,1})}{a_{1,1}}; \quad J_1(z) \sim a_{1,1}(z - j_{1,1}); \quad z \rightarrow j_{1,1}$$

- Exact condition for BH formation is the same (up to number of order 1) as the naive one, $\phi(b_{\max}) \sim 1$, $\Rightarrow \sigma = \pi b_{\max}^2$, namely

$$\sigma_{\text{tot}}(s) = \bar{K} \frac{\pi}{M_1^2} \ln^2[\sqrt{s} G_4 M_1 K], \quad K, \bar{K} \sim \mathcal{O}(1).$$

- More precisely, corrections: $\phi \sim 1 \Rightarrow$

$$\sigma_{\text{tot}}(s) = \frac{\pi}{M_1^2} \left[\ln(\sqrt{s} G_4 M_1 K) - \frac{1}{2} \ln(\ln(\sqrt{s} G_4 M_1 K)) \right]^2$$

- Exact map: $\phi(r, y = 0)$ to Heisenberg's DBI scalar. M_1 (first KK mode) $\leftrightarrow M$ of first glueball. Same $\phi \sim e^{-Mr} / \sqrt{Mr}$ behaviour.

- BH forming in collision of 2 gravitational shockwaves \leftrightarrow pion field "soliton" (fireball) formed in collision of 2 pion field shockwaves.
- Experiments: "Soft Pomeron" (until 2001, in PDG) \rightarrow small power law, $\sigma_{AB} \sim X_{AB}s^\epsilon$, $\sqrt{s} > 9\text{GeV}$. Later, used $Z_{AB} + B \log^2(s/s_0)$, extended down to $\sqrt{s} = 5\text{GeV}$, but with only *3 percent* better fit ($\chi^2/\text{d.o.f}$) \rightarrow simply $As^\epsilon \simeq A + \epsilon \log s + \epsilon^2/2 \log^2 s$.
- Find $B \sim 0.3\text{mb}$ instead of $\pi/m_\pi^2 \simeq 60\text{mb}$, but still $Z_{AB} \sim 18 - 65\text{mb}$ (now: *coincidence!*)
- Estimate E_F : is about a few TeV, such that

$$(s/s_0)^\epsilon = (\text{a few TeV}/9\text{GeV})^{0.093} \sim 60\text{mb}/X_{AB} = \frac{\pi/m_\pi^2}{X_{AB}}$$

4. Uniqueness and generalizations

- Other higher derivative actions? For

$$\mathcal{L} = -\frac{1}{2}(\partial_\mu\phi)^2 + \sum_{n \geq 2} C_n [(\partial_\mu\phi)^2]^n,$$

$\phi = A\sqrt{s}$ is not a solution:

$$\frac{A}{\sqrt{s}}(1 + \dots) + \left(\frac{A}{2\sqrt{s}} - \frac{A}{2\sqrt{s}} \right) \sum_{n \geq 2} (\dots) = 0.$$

- Only for DBI coefficients is a solution:

$$\frac{A}{\sqrt{s}} + \frac{\#}{1 - A^4 l^2} \left(\frac{A}{2\sqrt{s}} - \frac{A}{2\sqrt{s}} \right) = 0$$

is rewritten as an identity.

- Can add $V(\phi)$ instead of $m^2\phi^2/2$ without changing the solution as $s \rightarrow 0$.
- Can consider N scalar mesons ϕ^i with σ model in $d + 1$ dimensions, so DBI

$$\mathcal{L} = l^{-(d+1)} \left[h(\phi^i) - f(\phi) \sqrt{1 + l^{d+1} [g_{ij}(\phi^k) (\partial_\mu\phi^i) (\partial_\mu\phi^j) + 2V(\phi^i)]} \right]$$

- Find $\phi^a = A^a \sqrt{s}$, $\sum_a (A^a)^2 = l^{-4}$, and

$$\sigma_a = \frac{\pi}{m_a^2} \ln^2 \left(\frac{\sqrt{s}}{\langle k_0 \rangle} \right)$$

- General D-brane action with $g_{ij}(\phi) \simeq g(\phi) \delta_{ij} \Rightarrow$ *on the solution* $\phi = \phi(s)$ we have the same action as before \rightarrow same analysis.

- For AdS space ($m = 0$), $g(\phi) = \phi^{-4} \Rightarrow \phi(s) = 1/\sqrt{s}$ is exact solution. Also for general $g(\phi)$, $\int d\phi \sqrt{g(\phi)}$ is solution.

- Vector mesons:

$$\begin{aligned}\mathcal{L} &= l^{-4} \left[1 - \sqrt{\det(\eta_{ab} + l^4 \partial_a \phi \partial_b \phi + l^2 F_{ab}) + m^2 \phi^2 + M_V^2 A_a^2} \right] \\ &= l^{-4} \left[1 - \sqrt{1 + l^4 [(\partial_\mu \phi)^2 + m^2 \phi^2] + \frac{l^4}{2} F_{ab} F^{ab} - l^8 \left(\frac{1}{4} \tilde{F}_{ab} F^{ab} \right)^2 + M_V^2 A_a^2 + \dots} \right]\end{aligned}$$

- For just vector $A_a(r) = A_a e^{-M_V r}$. If also σ_V obtained in same way: $\langle k_0 \rangle = \sqrt{s} e^{-M_V b_{\max}} \Rightarrow$

$$\sigma_V = \pi b_{\max}^2 = \frac{\pi}{M_V^2} \ln^2 \frac{\sqrt{s}}{\langle k_0 \rangle}$$

5. Pion wavefunction

- In general, $\phi = \phi(s, r)$, $r = \sqrt{y^2 + z^2}$ (transverse)
- Near $s = 0$, r finite, find $\phi = A\sqrt{s} + s^n f(r)$, $n \geq 1$.
- Near $s = 0$, $r = 0$, find $\phi = A\sqrt{s} + s^{3/2} f(r) + s^{5/2} g(r)$.
- Delta function shockwave? Find that $\phi = \delta(x^+) \phi(r)$ unphysical. Can add by hand *source* to action (but doesn't explain source-free shockwave).
- But: only need $T_{++} \rightarrow \infty$ at $x^+ = 0$. Usually: $\delta(x^+)$ in $T_{++} \Rightarrow \delta(x^+)$ in ϕ . Now: slow blow-up of \mathcal{H} near $x^+ = 0$.

- Near $r \rightarrow \infty$, but a bit away, find *Bessel function* (solution to free wave eq.)

$$\phi = AK_0(mr) = A \frac{\pi i}{2} H_0^{(1)}(imr)$$

- At $r \rightarrow \infty$, find

$$\phi(r) \simeq A \sqrt{\frac{\pi}{2mr}} e^{-mr}$$

- At $r \rightarrow 0$, find

$$\phi(r) \simeq -A \ln \frac{mr}{2}$$

- Subleading corrections

$$\phi \simeq AK_0(mr) + \left(\frac{\pi}{2}\right)^{3/2} \frac{m^2 l^4 A^3}{4} \frac{e^{-3mr}}{(mr)^{3/2}}$$

6. Pion field source

- Source for pion shockwave: nucleons. How?

Vector BI

$$\mathcal{L} = \sqrt{1 + F - G^2} - 1$$
$$F \equiv \frac{1}{b^2}(\vec{B}^2 - \vec{E}^2); \quad G \equiv \frac{1}{b^2}(\vec{B} \cdot \vec{E}).$$

- Define also

$$\vec{H} \equiv b^2 \frac{\partial \mathcal{L}}{\partial \vec{B}}; \quad \vec{D} \equiv b^2 \frac{\partial \mathcal{L}}{\partial \vec{E}}$$

- Then find "Maxwell's equations"

$$\vec{\nabla} \times \vec{E} + \partial_0 \vec{B} = 0; \quad \vec{\nabla} \cdot \vec{B} = 0$$
$$\vec{\nabla} \times \vec{H} - \partial_0 \vec{D} = 0; \quad \vec{\nabla} \cdot \vec{D} = 0$$

- Then at $\vec{B} = \vec{H} = 0$, $\vec{\nabla} \cdot \vec{D} = 0$, which means

$$D_e = \frac{e}{r^2} \rightarrow \text{really } \vec{\nabla} \cdot \vec{D} = 4\pi e \delta^3(r).$$

- \vec{D} has *pointlike* source, though then

$$\rho = \frac{e}{2\pi r_0^3} \frac{1}{(r/r_0)(1 + (r/r_0)^4)^{3/2}}$$

is *spread out*.

- **Static scalar DBI**

$$\mathcal{L} = \sqrt{1 + \vec{F}^2}; \quad \vec{F} \equiv \vec{\nabla} \phi$$

- like the vector action for $\vec{B} = 0$, but different sign in r .

- Define $\vec{C} = \frac{\partial \mathcal{L}}{\partial \vec{F}}$, then $\vec{\nabla} \cdot \vec{C} = 0$, so find

$$C_r = \frac{e}{r^2} \rightarrow \text{really, } \vec{\nabla} \cdot \vec{C} = 4\pi e \delta^3(r).$$

- **DBI scalar shockwave**

$$\mathcal{L} = l^{-4} \left[1 - \sqrt{1 - 4l^4 s \left(\frac{d\phi}{ds} \right)^2} \right]$$

- Define analog of electric field

$$E_s = 2\sqrt{s} \frac{d\phi}{ds}$$

$$\mathcal{L} = l^{-4} \left[1 - \sqrt{1 - E_s^2} \right]$$

- Define also $D_s = \partial\mathcal{L}/\partial E_s$, so equation of motion is

$$\frac{d}{ds}(\sqrt{s}D_s) = 0 \Rightarrow D_s = \frac{A}{\sqrt{s}}, \quad \sigma > 0: \quad D_s = \frac{A}{\sqrt{s}}\theta(s).$$

- At $s \rightarrow 0$, $\phi(s) \simeq l^{-2}\sqrt{s}\theta(s)$, and $E_s = l^{-2}\theta(s)$.
- But D charge is delta function, whereas E charge is spread out:

$$\begin{aligned} \frac{d}{ds}(\sqrt{s}D_s) &= A\delta(s) \\ \rho &= \frac{d}{ds}(\sqrt{s}E_s) = \frac{l^4 A^2}{2\sqrt{s}(s + Al^4 A^2)^{3/2}}\theta(s). \end{aligned}$$

7. The cross section

- Corrections to leading behaviour of $\sigma_{\text{tot}}(s) \rightarrow$ in parallel for Heisenberg model and gravitational dual (shockwave with profile ϕ)

$$\phi(r) \propto \frac{e^{-mr}}{\sqrt{mr}} \quad \text{vs.} \quad \Phi(r) \propto \frac{e^{-M_1 r}}{\sqrt{M_1 r}}$$

- One finds $\sigma_{\text{tot}}(s) = \pi b_{\text{max}}^2(s)$, with

$$b_{\text{max}} \simeq \frac{1}{m_\pi} \ln \frac{\sqrt{\tilde{s}}}{\langle k_{0,\pi} \rangle} - \frac{1}{2m_\pi} \ln \left[\ln \frac{\sqrt{\tilde{s}}}{\langle k_{0,\pi} \rangle} \right]$$

- New regime? If $l^{-1} \sim \Lambda_{\text{QCD}}$, $m = m_\pi$, and $(lm)^2 \ll 1$, then *linear* wavefunction, and never get into nonlinear regime: $\phi(r) \simeq -A \ln \frac{mr}{2}$, leading to $(\sqrt{\tilde{s}} > \langle k_{0,\pi} \rangle)$

$$\sigma_{\text{tot}}(\tilde{s}) = \pi b_{\text{max}}(\tilde{s})^2 = \frac{\pi}{m_\pi^2} e^{-4 \frac{\langle k_{0,\pi} \rangle}{\sqrt{\tilde{s}}}}$$

- Black disk eikonal: $S = e^{i\delta}$, $\text{Im } \delta = \infty$ for $b \leq b_{\text{max}}(\tilde{s})$ and $\delta = 0$ for $b > b_{\text{max}}(\tilde{s})$. Then

$$\begin{aligned} \frac{1}{\tilde{s}} \mathcal{A}(\tilde{s}, t) &= -i \int d^2b e^{i\vec{q}\cdot\vec{b}} (e^{i\delta(b, \tilde{s})} - 1) \\ &= 2\pi i \frac{b_{\text{max}}(\tilde{s})}{\sqrt{t}} J_1(\sqrt{t} b_{\text{max}}(\tilde{s})) \end{aligned}$$

- Total cross section is

$$\frac{1}{\tilde{s}} \text{Im} \mathcal{A}_{\text{elastic}}(\tilde{s}, t = 0) = \sigma_{\text{tot}}(k_1, k_2 \rightarrow \text{anything}) = \pi b_{\text{max}}(\tilde{s})^2$$

- But differential cross section is

$$\frac{d\sigma}{dt} = \frac{|\mathcal{A}(\tilde{s}, t)|^2}{16\pi\tilde{s}(\tilde{s} - 4m^2)}.$$

- For black disk eikonal amplitude, find

$$\frac{\sigma_{\text{elastic}}}{\sigma_{\text{tot}}} = \frac{1}{4} \left[1 - \left(J_0(|b_{\text{max}}(\tilde{s})| \sqrt{\tilde{s} - 4m^2}) \right)^2 - \left(J_1(|b_{\text{max}}(\tilde{s})| \sqrt{\tilde{s} - 4m^2}) \right)^2 \right].$$

$$\simeq \frac{1}{4} \left[1 - \frac{2}{\pi b_{\text{max}}(\tilde{s}) \sqrt{\tilde{s} - 4m_N^2}} \right]$$

- Translating into QCD, with Heisenberg's model,

$$\frac{\sigma_{\text{elastic}}}{\sigma_{\text{tot}}} \simeq \frac{1}{4} \left[1 - \frac{2m_\pi}{\pi \ln(\tilde{s}/s_0) \sqrt{\tilde{s} - 4m_N^2}} \right]$$

- Compares well with experimental results.

8. Holographic models

- One model: hard-wall (Polchinski-Strassler)
- But any gravity dual can be approximated by it. Details of IR only matter for $E_F \rightarrow$ energy at which Froissart behaviour onsets. Otherwise general.
- Note that ∞ nr. of higher derivative terms in DBI \rightarrow chiral perturbation theory \leftrightarrow geometry of D-brane embedding
- Other model: Sakai-Sugimoto type.

- Consider a double Wick rotation of a nonextremal D4 background, with a $D8 - \bar{D}8$ pair forming a U-shape.

$$ds^2 = \left(\frac{u}{R_{D4}}\right)^{3/2} \left[-dt^2 + \delta_{ij} dx^i dx^j + f(u) dx_4^2 \right] + \left(\frac{R_{D4}}{u}\right)^{3/2} \left[\frac{du^2}{f(u)} + u^2 d\Omega_4^2 \right]$$

$$F_4 = \frac{2\pi N_c}{V_4} \epsilon_4 \quad , \quad e^\phi = g_s \left(\frac{u}{R_{D4}}\right)^{3/4} \quad , \quad R_{D4}^3 = \pi g_s N_c l_s^3 \quad , \quad f(u) = 1 - \left(\frac{u_\Lambda}{u}\right)^3$$

- The D8-brane action, integrated over transverse coordinates, is

$$S_{DBI} = \tilde{T}_8 \int dt d^3x dx_4 \phi^4 \sqrt{f(\phi) + \left(\frac{R_{D4}}{\phi}\right)^3 \left[\partial_\mu \phi \partial^\mu \phi + \frac{1}{f(\phi)} (\partial_{x_4} \phi)^2 \right]}$$

- Expand $\phi(x^\mu, x^4)$ in KK modes and integrate over massive modes (with wavefunction $\zeta_n(x_4)$) \rightarrow leave an action for the lowest one, assumed to be of the type

$$S_{DBI} = \tilde{T}_8 \int dt d^3x \phi^4 \sqrt{f(\phi) + \left(\frac{R_{D4}}{\phi}\right)^3 \left[\partial_\mu \phi \partial^\mu \phi + m^2 \phi^2 \right]}.$$

- Protons \rightarrow can be thought of as wrapped D4-branes on the flavor brane giving an instanton-like structure, and action like the one above.

9. Relativistic hydrodynamics and DBI

- We want to obtain a hydro description *from* the Heisenberg one via the shock of the shockwave solution.
- Belinfante energy-momentum tensor: $T_{00} \rightarrow \infty$ at shock.
- Expanding DBI, find coefficient of $\phi^2(\partial_\mu\phi)^2$ and identify with $f_\pi^{-2}/6$, giving

$$\beta \equiv l^2 = \frac{\sqrt{2}}{\sqrt{3}m_\pi f_\pi} \simeq \frac{1}{(126\text{MeV})^2}$$

- Relativistic hydro: expansion in ∂u ($a_\mu = u^\rho \nabla_\rho u_\mu$)

$$T_{\mu\nu} = \rho u_\mu u_\nu + P(g_{\mu\nu} + u_\mu u_\nu) + \pi_{\mu\nu}$$

$$\pi_{\mu\nu} = -2\eta \left[\frac{\nabla_\mu u_\nu + \nabla_\nu u_\mu}{2} + \frac{a_\mu u_\nu + a_\nu u_\mu}{2} - \frac{1}{3}(\nabla^\rho u_\rho)(g_{\mu\nu} + u_\mu u_\nu) \right]$$

$$-\zeta(\nabla^\mu u_\mu)(g_{\mu\nu} + u_\mu u_\nu) + \dots$$

- Eqs. of motion (relativistic generalization of Navier-Stokes) are $\partial^\mu T_{\mu\nu} = 0$.

- Obtain (in non-rel. limit: Navier-Stokes)

$$\begin{aligned}
 & u_\mu u_\nu \partial^\mu (\rho + P) + (\rho + P) (\partial^\mu u_\mu) u_\nu + (\rho + P) u_\mu \partial^\mu u_\nu + \partial_\nu P \\
 &= \eta \partial^\mu (\partial_\mu u_\nu + \partial_\nu u_\mu) + \left(\zeta - \frac{2\eta}{3} \right) \partial^\mu [(\partial^\rho u_\rho) (g_{\mu\nu} + u_\mu u_\nu)]
 \end{aligned}$$

- Conformal fluids: If $T^\mu{}_\mu = 0 \Rightarrow \rho = 3P, \zeta = 0$.
- But for $d \neq 2$, conf. inv. $\not\Rightarrow T^\mu{}_\mu = 0$. E.g. massless scalar:

$$S = \int d^4x \left[-\frac{1}{2} (\partial_\mu \phi)^2 \right] \Rightarrow T^{\mu\nu} = \frac{2-d}{2} (\partial_\rho \phi)^2.$$

- Use Noether ambiguity $T^{\mu\nu} \rightarrow T^{\mu\nu} + \partial_\lambda J^{\mu\nu\lambda}$.

• Now

$$T_{\mu\nu}^I = T_{\mu\nu}^B - \frac{1}{6}(\partial_\mu\partial_\nu - g_{\mu\nu}\partial^2)\phi^2$$

is traceless on-shell ($\square\phi = 0$)

$$T^{I\mu}{}_\mu = T^{B\mu}{}_\mu + (\partial_\mu\phi)^2 + \phi\partial^2\phi$$

10. DBI scalar hydrodynamics

- Heisenberg model worked well in sQGP regime ($s \rightarrow \infty$), so hydro description should *follow* from it.
- What is u^μ ? $u^\mu = k^\mu/m \propto \partial^\mu \phi$:

$$u^\mu = \frac{\partial^\mu \phi}{\sqrt{-(\partial_\mu \phi)^2}}$$

- Consistent with usual scalar field $\int d^2x [-(\partial\phi)^2/2 - V]$. Then $u^\mu = (1, 0, 0, 0)$ and

$$\rho_\phi = \frac{\dot{\phi}^2}{2} + V; \quad p_\phi = \frac{\dot{\phi}^2}{2} - V$$

- Free massless scalar: use improved $T_{\mu\nu}$. Then, identify with $T_{\mu\nu}$ expansion in $u^\mu \Rightarrow$

$$\rho = -\frac{1}{2}(\partial_\mu\phi)^2; \quad P = -\frac{1}{6}(\partial_\mu\phi)^2 \Rightarrow P = \rho/3$$

- Identifying the coefficient of the $\partial_\mu u_\nu$ term, for the free massless scalar, we find

$$\zeta = 0; \quad 2\eta = \frac{\phi}{3}\sqrt{-(\partial_\lambda\phi)^2} \rightarrow \infty \text{ as } s \rightarrow 0.$$

- However, the important object is the ratio

$$\frac{\eta}{s} = \frac{\eta T}{\rho + P} = \frac{\phi T}{4\sqrt{-(\partial_\lambda\phi)^2}} \rightarrow 0; \text{ as } s \rightarrow 0.$$

- Non-viscous fluid.

- Ideal DBI hydro: generalize to DBI.

- Improved tensor that is traceless as $\beta \rightarrow 0$ is

$$T_{\mu\nu}^I = \frac{\frac{2}{3}(\partial_\mu\phi\partial_\nu\phi - g_{\mu\nu}(\partial_\rho\phi)^2) - g_{\mu\nu}\beta^{-2}}{\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} + \frac{g_{\mu\nu}}{\beta^2} - \frac{\phi}{3\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} \left(\partial_\mu\partial_\nu\phi - \frac{\beta^2\partial_\nu\phi(\partial_\mu\partial^\rho\phi)\partial_\rho\phi}{1 + \beta^2(\partial_\lambda\phi)^2} \right).$$

- The velocity is the same $u^\mu = \partial^\mu\phi/\sqrt{-(\partial_\lambda\phi)^2} \rightarrow \beta^2\partial^\mu\phi$ as $s \rightarrow 0$.

- Find

$$\rho = \frac{1}{\beta^2\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} - \frac{1}{\beta^2}$$

$$P = \frac{1}{\beta^2}(1 - \sqrt{1 + \beta^2(\partial_\lambda\phi)^2}) + \frac{(\partial_\rho\phi)^2}{3\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}}$$

- But that gives $(3P + \rho)/\rho \rightarrow 0$ ($P \simeq -\rho/3$). Wrong.

- With another improved tensor, we get tracelessness near the shock:

$$T_{\mu\nu}^I = \frac{\frac{4}{3}(\partial_\mu\phi\partial_\nu\phi - g_{\mu\nu}(\partial_\rho\phi)^2) - g_{\mu\nu}\beta^{-2}}{\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} + \frac{g_{\mu\nu}}{\beta^2} + \frac{\phi}{3\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} \left(\partial_\mu\partial_\nu\phi - \frac{\beta^2\partial_\nu\phi(\partial_\mu\partial^\rho\phi)\partial_\rho\phi}{1 + \beta^2(\partial_\lambda\phi)^2} \right).$$

- Find

$$\rho = \frac{1}{\beta^2\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} - \frac{1}{\beta^2}$$

$$P = \frac{1}{\beta^2}(1 - \sqrt{1 + \beta^2(\partial_\lambda\phi)^2}) - \frac{(\partial_\rho\phi)^2}{3\sqrt{1 + \beta^2(\partial_\lambda\phi)^2}} \simeq \frac{\rho}{3} \rightarrow \infty$$

- Then we find viscosity:

$$\eta = \frac{\phi \sqrt{-(\partial_\lambda \phi)^2}}{3 \sqrt{1 + \beta^2 (\partial_\mu \phi)^2}}$$

- But near the shock

$$\frac{\eta}{\rho + P} = \frac{\phi}{4 \sqrt{-(\partial_\lambda \phi)^2}} \rightarrow 0$$

- So DBI is "ideal hydro"
- Can add mass m and conclusions don't change.

- Viscous DBI: add term inside the square root of DBI

- Modified DBI is

$$S = -\beta^{-2} \int d^4x \left[\sqrt{1 + \beta^2 [(\partial\phi)^2 + m^2\phi^2]} + \alpha \left[\partial^2\phi - \frac{(\partial_\mu\phi)(\partial^\mu\partial^\rho\phi)(\partial_\rho\phi)}{(\partial_\lambda\phi)^2} \right] - 1 \right]$$

- On the $\phi = \phi(s)$ solution, it becomes

$$S = -\beta^{-2} \int d^4x \left[\sqrt{1 + \beta^2 \left(-4s \left(\frac{d\phi}{ds} \right)^2 + m^2\phi^2 \right)} + \alpha \left(-2 \frac{d\phi}{ds} \right) - 1 \right]$$

- Shear viscosity is then

$$2\eta = \frac{\alpha \sqrt{-(\partial_\lambda\phi)^2}}{\beta^2 \sqrt{1 + \beta^2 [(\partial\phi)^2 + m^2\phi^2]} + \alpha \left[\partial^2\phi - \frac{(\partial_\mu\phi)(\partial^\mu\partial^\rho\phi)(\partial_\rho\phi)}{(\partial_\lambda\phi)^2} \right]} \rightarrow \infty$$

and bulks is $\zeta = 2\eta/3$.

- Obs. Relativistic Navier-Stokes is *satisfied*.

$$\frac{\eta}{s} = \frac{\eta T}{\rho + P} = T \frac{\alpha}{\beta^2 \sqrt{-(\partial_\lambda \phi)^2}} \rightarrow T \frac{\alpha}{\beta}$$

is finite!

- Fix parameters from experiment. On-shell,

$$\sim \frac{\alpha^2}{\beta^2} (\partial^2 \phi)^2 \sim \frac{\alpha^2 m_\pi^4}{\beta^2} \phi^2 \leftrightarrow \sim m^2 \phi^2.$$

giving

$$\alpha \sim \frac{\beta}{m_\pi} = \frac{\sqrt{2}}{\sqrt{3} f_\pi m_\pi^2}.$$

- Then also

$$T \sim m_\pi \frac{\eta}{s} \sim m_\pi \left(\times \frac{1}{4\pi} \right)$$

- If $T \simeq 4m_\pi/\pi \Rightarrow$ conjectured in HN 2015, then

$$\alpha \simeq \frac{\beta}{16m_\pi}$$

- Viscosity regulates shock. Solution $\phi \simeq \sqrt{s}/\beta$ not valid anymore. We have instead

$$\phi(s) \simeq A + BsCs^2 + \dots, \quad \sqrt{s} \ll \frac{3\alpha}{4\beta} \propto \frac{1}{m_\pi}$$

- and the solution $\phi(s) \simeq \sqrt{s}/\beta$ valid for (if \exists small nrs. like 1/16 above)

$$\frac{3\alpha}{4\beta} \ll \sqrt{s} \ll \frac{1}{m_\pi}$$

- Extra term regulates behaviour: denominator is finite in $\mathcal{H} \rightarrow$ finite energy density at shock.

Conclusions

- Heisenberg model describes well $s \rightarrow \infty$ regime: Froissart bound like in AdS/CFT.
- Is unique, but can be generalized by including $V(\phi)$, several scalar mesons, vector mesons, extra functions.
- For near-asymptotic wavefunction, obtain corrections to $\sigma_{\text{tot}}(s)$ away from Froissart.
- The shockwave has dual interpretation: delta function at shock $x^+ = 0$, and spread out, similar to vector DBI.
- We can also find $\sigma_{\text{elastic}}/\sigma_{\text{tot}}(s)$ under certain assumptions
- Ideal hydrodynamics is derived from DBI.
- Can introduce finite η as a regulator extra term in DBI, and then $\zeta = 2\eta/3$.